Chapter 3. Quantum Mechanics Postulates, Principles, and Mathematic Formalism § 3.1. Postulates of Quantum Mechanics

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- 1. QM State of a system is defined as the undisturbed motion of the system. The state is about the Status of a motion of the physical system. Each undisturbed motion with information about energy, momentum, coordinates, angular momentum etc forms a state of the physical system at an instant time. The change of the state with time, the change of motion, or the disturbed motion belongs to the problem of time evolution of the state.
- * In QM mathematical formalism, the state of any physical system at an instant time is represented by a state vector (4(t)) in an abstract state space that is formed by all possible states of the physical system.
- 2. QM Physical quantities / dynamic variables / observables are the physical quantities that Can be measured or observed. These observables are represented by abstract linear operators A.
- 3. QM measurements are represented by a linear operator A acting on a state vector 14>, i.e., A 14>. The only possible result of the measurement is one of the eigenvalues of the Corresponding observable A.

4. QM Eigenvalue Equation of a linear operator is defined as 14> = λ 14>
where 14> is the eigenstate vector of the operator Â, λ is a complex number, the eigenvalue of Â.

- 5. QM state vectors and linear operators obey the followy rules,
 - (1) Ket vector 1> and Bra vector <1 are "Hermitian Conjugate" With each other: $<1=1>^{*}$ (14> = <41*)
 - (2) The scalar product of two ket vectors $|\Psi_i\rangle$ and $|\Psi_2\rangle$ is given by the product of one ket vector and the conjugate of the other ket vector: $|\Psi_i\rangle^* \cdot |\Psi_2\rangle = \langle \Psi_i|\Psi_2\rangle$ i.e., the bra vector corresponding to one ket vector, times the other ket vector.
 - (3) When taky conjugate of state vector, operator, and constant, we take the conjugate of each and then reverse the order, e.g., $(\lambda \hat{A} | \Psi \rangle)^* = (\langle \Psi | \hat{A}^* \lambda^* \rangle)$

(4) If $\hat{A} = A^*$, then \hat{A} is a Hermitian operator.

(5) Other linear operation rules:
$$\hat{A}|\Psi\rangle = |\Psi'\rangle$$
,
 $(\hat{A}\hat{B})|\Psi\rangle = \hat{A}(\hat{B}|\Psi\rangle)$
 $\langle \Psi_1|(\hat{A}|\Psi_2\rangle) = (\langle \Psi_1|\hat{A}\rangle)|\Psi_2\rangle$
 $\hat{A}(\lambda_1|\Psi_1\rangle + \lambda_2|\Psi_2\rangle) = \lambda_1\hat{A}|\Psi_1\rangle + \lambda_2\hat{A}|\Psi_2\rangle$

(6) In general, $\hat{A}\hat{B} \neq \hat{B}\hat{A}$. The commutator is $[\hat{A}, \hat{B}] = \hat{A}\hat{B} - \hat{B}\hat{A}$. \$3.2. Principle of Superposition of States

6. QM principle: of Superposition of states requires that U. Any state of a system can be considered as a superposition of two or more other states of the same system, and indeed in an infinite number of ways. 2). The superposition of two or more states of a system forms a new state of the same system. In QM, if a physical system has a complete orthonormal set of eigenstates { [Un> } with discrete eigenvalues { an } of a linear operator \hat{A} : $\hat{A} | u_n \rangle = a_n | u_n \rangle$, $\langle u_m | u_n \rangle = \delta_{mn}$, then any state of the system 14> can be expressed in terms of the complete set { 1 un> }. $| \psi \rangle = \sum_{n} (c_n | u_n \rangle)$, where $c_n = \langle u_n | \psi \rangle$. If a system has a complete orthonormal set of eigenstates { | Wa>} with continuous eigenvalues { a linear operator À: Â | Wa> = ~ | Wa>, < War | Wa> = dia-a') then any state of this system 142 can be expressed in terms of the complete set { | Wa>}: $|\psi\rangle = \int C(\alpha) |\psi\rangle d\alpha$, where $C(\alpha) = \langle \psi | \psi \rangle$. These are the mathematical expression of QM principle of Superposition of states.

- 7. QM principle of spectral decomposition is to describe the probability of obtaining a specific eigenvalue of an observable operator when making measurements of a system in certain state.
- (1) If an observable has eigenvalues {an}(discrete) associated with its eigenstates {1 kn >}, i.e., Â | kn > = an | kn >. When the physical quantity is measured on the system in the state 14>, the probability of obtaining a non-degenerate

eigenvalue
$$a_n$$
 of \hat{A} is given by $|\langle u_n | \psi \rangle|^2 = |C_n|^2$

$$P(an) = \frac{|\langle un| \psi \rangle|}{\langle \psi| \psi \rangle} = \frac{|\langle un| \psi \rangle|}{\langle \psi| \psi \rangle}$$

where $|\psi\rangle = \sum_{n} C_{n}|u_{n}\rangle$, and $C_{n} = \leq u_{n}|\psi\rangle$ is called the probability amplitude.

(2) In general, if an is degenerate, the probability is given
by
$$P(a_n) = \frac{\frac{3n}{2} |\langle u_n^i | u \rangle|^2}{\langle u_n^i | u \rangle} = \frac{\frac{3n}{2} |\langle u_n^i |^2}{\langle u_n^i | u \rangle}$$

(3) If an observable A has continuous eigenvalues {X} associated
With eigenstates {|Wx>}, i.e., A |Wx> = x |Wx>,
When the physical quantity A is measured on the system
in the state |W>, the probability of obtaining a non-
degenerate eigenvalue between a and x+dx is given by

$$dP(x) = \frac{|\langle Wx| W >|^2}{\langle V| W >} dx = \frac{|C(x)|^2}{\langle V| W >} dx$$

Where $|W> = \int C(x) |Wx> dx$, and $|C(x)|^2 = \frac{|\langle Wx| W >|^2}{\langle V| W >}$ is probability
 $density$.

- 8. Projector operator and two useful relations:
 - (1) $|4\rangle < 9|$ becomes an operator, because when it applies to an arbitrary ket vector, it gives another ket vector. Take an arbitrary ket vector and apply $|4\rangle < 9|$ to it. $(|4\rangle < 9|) |A\rangle = |4\rangle < 9|A\rangle$

Since <91A> is the scalar product so a complex number, thus, above equation gives another ket vector. Therefore, 14><91 is an operator.

(2) Let 14> be a normalized ket vector, i.e., <414>=1. The operator $P_{4} = |4\rangle < |4|$ is a projector operator that projects any arbitrary ket vector 19> onto the ket vector $|\psi\rangle$: $P_{\psi}|\varphi\rangle = |\psi\rangle \langle \psi|\varphi\rangle$ This is further confirmed by the fact $P_4^2 = P_4$, i.e., projecting twice in succession onto a given vector is equivalent to projecting a single time. $P_{\psi}^{2} = P_{\psi}P_{\psi} = |\psi\rangle < \psi|\psi\rangle < \psi| = |\psi\rangle < \psi| = P_{\psi}$ (3) For a complete orthonormal set of eigenstates { | Un> } We have the projector operator $p = \sum_{m} |U_m\rangle < U_m| = 1$. In a continuous eigenvalue case { I Wa>}, $P = \int |Wa \rangle \langle Wa | da = 1$ This can be derived as below.

(4) For a complete orthornormal set of eigenstates
$$\{|\mathcal{U}_{n}\rangle\}$$

any state $|\Psi\rangle$ can be expressed as: $\{\mathcal{U}_{n}|\mathcal{U}_{n}\rangle = \overline{\delta_{nn}}\}$
 $|\Psi\rangle = \sum_{n} C_{n}|\mathcal{U}_{n}\rangle$.
where $C_{n} = \langle \mathcal{U}_{n}|\Psi\rangle$. Substitute C_{n} into above equation.
Then, $|\Psi\rangle = \sum_{n} C_{n}|\mathcal{U}_{n}\rangle$
 $= \sum_{n} \langle \mathcal{U}_{n}|\Psi\rangle |\mathcal{U}_{n}\rangle$
As $\langle \mathcal{U}_{n}|\Psi\rangle$ is a scalar product, i.e., a complex number,
it can switch order with $|\mathcal{U}_{n}\rangle$. Thus,
 $|\Psi\rangle = \sum_{n} |\mathcal{U}_{n}\rangle \langle \mathcal{U}_{n}|\Psi\rangle = (\sum_{n} |\mathcal{U}_{n}\rangle \langle \mathcal{U}_{n}|)|\Psi\rangle$.
Because this equation is true for any arbitrary $|\Psi\rangle$,
the only conclusion is $\sum_{n} |\mathcal{U}_{n}\rangle \langle \mathcal{U}_{n}| = 1$.
(5) Similarly, for a continuous case $|\{W_{n}\rangle\}$ and
 $\langle W_{n'}|W_{n}\rangle = \overline{\delta}(\alpha - \alpha')$. Any state $|\Psi\rangle$ is expressed
as $|\Psi\rangle = \int C(\alpha) |W_{n}\rangle d\alpha$.
Where $C(\alpha) = \langle W_{n}|\Psi\rangle$. Substitute $C(\alpha)$ into above equation.
 $|\Psi\rangle = \int \langle W_{n}|\Psi\rangle |W_{n}\rangle d\alpha$
 $= (\int |W_{n}\rangle \langle W_{n}| d\alpha) |\Psi\rangle$.
Since this is true for any $|\Psi\rangle$, we conclude
 $\int |W_{n}\rangle \langle W_{n}| d\alpha = 1$

§3.3 Quantum Mechanics Measurements

9. Mean value of an observable \hat{A} from many times of measurements on a system in the state $|\psi\rangle$ is given by $\overline{A} = \langle \psi | \hat{A} | \psi \rangle$.

To understand this, $\langle 4| = |4\rangle^*$, i.e., $\langle 4|$ is the Conjugate of $|4\rangle$. $\hat{A}|4\rangle$ becomes another ket vector. So above equation implicates that the mean of \hat{A} should be calculated by using $|4\rangle^*$ to take a scalar product with $(\hat{A}|4\rangle)$ ket vector.

This mean-value equation is true for all possible cases, regardless discrete or continuous eigenvalues. It is

a natural conclusion from the principle of spectral decomposition. À has a complete orthonormal eigenstate set {IUn>} or {IWa>}. À IUn>=anIUn>, À IWa>=a/Wa>. Any state IU> can be expressed as

$$|\psi\rangle = \sum_{n} C_{n} |\mathcal{U}_{n}\rangle$$
 or $|\psi\rangle = \int C(\alpha) |\mathcal{W}_{n}\rangle d\alpha$.

where $C_n = \langle u_n | \psi \rangle$, or $C(\alpha) = \langle w_n | \psi \rangle$.

The probability of obtaining an result is $P(a_n) = |Kun||k\rangle|^2$, and for $\alpha = |\alpha| + d\alpha$, $dP(\alpha) = |Kun||k\rangle|^2 d\alpha$. Thus, the mean of \hat{A} should be calculated as

$$\overline{A} = \sum_{n} [P(a_n) a_n] \quad \text{or } \overline{A} = \int \alpha \, dP(\alpha)$$
$$= \sum_{n} [|\langle u_n|\psi \rangle|^2 a_n] \quad = \int |\langle u_n|\psi \rangle|^2 \alpha \, d\alpha.$$

Now let us verify whether
$$\langle \psi | \hat{A} | \psi \rangle$$
 is equivalent to
the above two means of \hat{A} .
In discrete eigenvalue case, $|\psi\rangle^{*} = \sum_{n} C_{n}^{*} \langle u_{n}|$
 $\langle \psi | \hat{A} | \psi \rangle = (\sum_{n} C_{n}^{*} \langle u_{n}|) \hat{A} (\sum_{m} C_{n} | u_{m} \rangle)$
 $= \sum_{n} \sum_{m} (C_{n}^{*} C_{n} \langle u_{n} | \hat{A} | u_{m} \rangle)$
 $= \sum_{n} \sum_{m} (C_{n}^{*} C_{n} \alpha_{m} \langle u_{n} | u_{m} \rangle)$
 $= \sum_{n} \sum_{m} (C_{n}^{*} C_{n} \alpha_{m} \langle u_{n} | u_{m} \rangle)$
 $= \sum_{n} \sum_{n} (C_{n}^{*} C_{n} \alpha_{m} \langle u_{n} | u_{m} \rangle)$
 $= \sum_{n} \sum_{n} (C_{n}^{*} C_{n} \alpha_{m} \delta_{nm})$
 $= \sum_{n} |C_{n}|^{2} \alpha_{n}$
 $= \sum_{n} |C_{n}|^{2} \alpha_{n}$
 $= \sum_{n} |C_{n}|^{2} \alpha_{n}$ - Equivalent !
In continuous eigenvalue case, $|\psi\rangle^{*} = \int C^{*}(\alpha) \langle u_{m}| d\alpha$
 $\langle \psi | \hat{A} | \psi \rangle = (\int C^{*}(\alpha) \langle u_{m}| d\alpha) \hat{A} (\int C(\alpha^{*}) | u_{m} \rangle d\alpha)$
 $= \iint C^{*}(\alpha) C(\alpha^{*}) \langle u_{m}| \dot{A} | u_{m} \rangle d\alpha d\alpha^{*}$
 $= \iint C^{*}(\alpha) C(\alpha^{*}) \alpha^{*} \delta_{m} d\alpha^{*}$
 $= \iint C^{*}(\alpha) C(\alpha^{*}) \alpha^{*} \delta_{m} d\alpha^{*}$
 $= \int |C_{m}|^{2} \alpha d\alpha$
 $= \int |C_{m}|^{2} \alpha d\alpha$, $-E_{guivalent}!$

10. QM reduction of the state: assume that we make measurement of \hat{A} on a system in the state $|\Psi\rangle$. Let us first consider the case where the measurement of \hat{A} yields a simple eigenvalue Ω_n of the observable \hat{A} . QM postulates the state of the system immediately after the measurement is the eigenstate $|U_n\rangle$ associated with Ω_n : $|\Psi\rangle \stackrel{(\Omega_n)}{\Longrightarrow} |U_n\rangle$

If we perform a second measurement of \hat{A} immediately after the first one (that is, before the system has had time to evolve), We shall always find the same result an, since the state of the system immediately before the second measurement is |un> (the eigenstate), and no longer 14>.

When the eigenvalue an given by the measurement is degenerate, the reduction of the state can be generalized as follows. If the expansion of the state IW> immediately before the measurement is written as $|W> = \sum_{n} \left(\sum_{i=1}^{g_n} C_n^i |U_n^i > \right)$, where g_n is the degeneracy factor of eigenvalue a_n , the reduction of the state becomes $|W> \frac{(a_n)}{\sqrt{\sum_{i=1}^{g_n} |C_n^i|^2}} \sum_{i=1}^{g_n} C_n^i |U_n^i > \sum_{$

11. QM Probability Amplitude and Interference Effect, Let 14, > and 14=> be two orthogonal normalized states . $\langle \Psi_1 | \Psi_2 \rangle = 0, \quad \langle \Psi_1 | \Psi_1 \rangle = \langle \Psi_2 | \Psi_2 \rangle = 1.$ For a given observable Â, {IUn>} are complete orthonormal eigenstates corresponding to the eigenvalues {an}. Alun>=an/Un>, <um/Un>=Jmn. * If the system is in the state 14,>, the probability P. (an) of finding an when A is measured on the system is given by $P_4(a_n) = |\langle u_n | \psi_i \rangle|^2$ * If the system is in the state $|\Psi_2\rangle$, the probability $P_2(a_n)$ of finding an when A is measured on the system is given by $P_{z}(a_{n}) = |\langle u_{n}| \Psi_{z} \rangle|^{2}$ * Now consider a normalized state 14> which is a linear superposition of $|\psi_1\rangle$ and $|\psi_2\rangle$: $|\psi\rangle = \lambda_1 |\psi_1\rangle + \lambda_2 |\psi_2\rangle$ $|\lambda_1|^2 + |\lambda_2|^2 = 1.$ The probability P(an) of finding the eigenvalue an when the observable à is measured on the system in the state 14> is $P(a_n) = |\langle u_n | \psi \rangle|$ Substituting 14> expression into Pcans and further deriving it: $P(a_n) = |\langle u_n | \psi \rangle|^2 = |\lambda_1 \langle u_n | \psi_1 \rangle + \lambda_2 \langle u_n | \psi_2 \rangle|^2$ = $(\lambda_1, U_n | \psi_1 > + \lambda_2 < U_n | \psi_2 >) (\lambda_1^* < U_n | \psi_1 > + \lambda_2^* < U_n | \psi_2 >)$ $= |\lambda_{i}|^{2} |\langle u_{n} | \Psi_{i} \rangle|^{2} + |\lambda_{z}|^{2} |\langle u_{n} | \Psi_{z} \rangle|^{2}$ + $\lambda_1 \lambda_2^* < u_n | \psi_1 > < u_n | \psi_2 > + \lambda_1^* \lambda_2 < u_n | \psi_1 > < u_n | \psi_2 >$ $= |\lambda_1|^2 P_1(a_n) + |\lambda_2|^2 P_2(a_n) + 2Re(\lambda_1\lambda_2^* < u_n|\psi_1 > < u_n|\psi_2 >)$ Thus, the Cross term shows the interference effect , Since P(a), P. (an) and Pz (an) are the probabilities, <un 14>, <un 14>, and < Un 142> are called the probability amplitudes. The interference effect is explained by the superposition of the probality amplitudes, $\langle u_n | \psi \rangle = \lambda_1 \langle u_n | \psi_1 \rangle + \lambda_2 \langle u_n | \psi_2 \rangle$

In last lecture of QM basic,
- State: |U>
- Observable: Â
- Measurement |U>
- IF |U> = C1 |U> + C2 |U2> + C3 |U3>, Superposition of states,
- any single measurementer gives either Q1, Q2, or Q3.
i.e., one of the eigen values.
This is the reduction/collapse of state during measurements.
- Mean value of an observable:

$$\overline{A} = \langle U | \widehat{A} | U \rangle$$
 $| \Psi \rangle = \sum_{n=1}^{\infty} C_n | U_n \rangle$

$$4 = \langle U | A | U \rangle \qquad | \Psi \rangle = \sum_{n} C_{n} | U_{n}$$
$$= \sum_{n} (| \langle U_{n} | \Psi \rangle|^{2} a_{n})$$

Probability of an occurrence

12. QM Commutation Relation.
Consider coordinate
$$\hat{\chi}$$
 and momentum \hat{p}_{χ} , $\hat{p}_{\chi} = -i\hbar \frac{\partial}{\partial \chi}$
in the $\{1\times\}$ representation - For any arbitrary state $|\Psi\rangle$ in
the $\{1\times\}$ representation $\Psi(x)$,
 $\hat{p}_{\chi} \hat{\chi} \Psi(x) = \hat{p}_{\chi} [\hat{\chi} \Psi(x)]$
 $= -i\hbar \frac{\partial}{\partial \chi} \chi \Psi(x) - i\hbar \times \frac{\partial}{\partial \chi} \Psi(x)$
 $= -i\hbar \frac{\partial}{\partial \chi} \chi \Psi(x) - i\hbar \times \frac{\partial}{\partial \chi} \Psi(x)$
 $= -i\hbar \Psi(x) - \chi \cdot i\hbar \frac{\partial}{\partial \chi} \Psi(x)$
 $= -i\hbar \Psi(x) - \chi \cdot i\hbar \frac{\partial}{\partial \chi} \Psi(x)$
 $= (i\hbar + \hat{\chi} \hat{p}_{\chi}) \Psi(x)$.
Since $\Psi(x)$ is arbitrary, we conclude that
 $\hat{\chi} \hat{p}_{\chi} - \hat{p}_{\chi} \hat{\chi} = i\hbar$
Define $[\hat{\chi}, \hat{p}_{\chi}] = \hat{\chi} \hat{p}_{\chi} - \hat{p}_{\chi} \hat{\chi}$,
 $\therefore [\hat{\chi}, \hat{p}_{\chi}] = i\hbar$. But $[\hat{\chi}, \hat{p}_{\chi}] = D$
Similarly, $[\hat{\psi}, \hat{p}_{\chi}] = i\hbar$
 $\sum [\hat{h}_{\chi}, \hat{p}_{\chi}] = i\hbar \delta_{\chi}$, or $[\hat{A}, \hat{B}] = \hat{A}\hat{B} - \hat{B}\hat{A}$
is called the commutation velation.

§3.4. Principle of Uncertainty – Indeterminacy

0. Expression of Heisenberg's Principle of Uncertainty

It is worth to point out that the principle of uncertainty originates from the wave-particle duality of matter. It is a nature's law, not due to measurement technologies.

1. Interpretation of Uncertainty Principle: It is impossible to determine the position and momentum of a particle to an arbitrary degree of accuracy.

Examples to show the effects of $h \neq 0$ in macro-world and in micro-world

And demonstrate why a motion orbit is meaningful only in macro-world.

2. Derivation of Uncertainty Principle from Commutation Relation

This is a good reference for students but not to be discussed in class (due to time constraint). This derivation shows that the uncertainty relation can be derived exactly from quantum mechanics' commutation relation. The principle of uncertainty is fundamental to quantum mechanics, but not an approximation.

3. Generalization of Heisenberg's Principle of Uncertainty

$$\Delta A \cdot \Delta B \ge \frac{1}{2} |< [\hat{A}, \hat{B}] > |$$

Derive $\Delta E \cdot \Delta t \ge \frac{\hbar}{2}$ from $\Delta p \cdot \Delta x \ge \frac{\hbar}{2}$

4. The Intrinsic Nature of Uncertainty Principle is the Wave-Particle Duality of Matter (all material particles, photons and electromagnetic waves). It is a law of the nature, not due to measurement technologies or imperfection.

Give an example of electron diffraction to demonstrate

- 1) Why the uncertainty is on the order of h Planck constant.
- 2) The unequal fringe heights in the double-slit experiment are due to the modulation of interference fringes by diffraction.
- 3) Diffraction's nature is the multiple-beam interference.

5. Applications of Principle of Uncertainty

This is to demonstrate the values of this principle in solving real problems, in which experimental facts cannot be explained well by classical physics.

- 1) The minimum kinetic energy of bound particle
- 2) Electrons inside an atom cannot fall in the nucleus
- 3) Natural linewidth of spectral lines

\$3.4 Principle of Uncertainty - Indeterminacy

13. QM Principle of Uncertainty is a fundamental principle in Quantum Mechanics. It puts a fundamental limit on the accuracy of simultaneously determining the numerical values of two non-Commuting observables. When these two non-Commuty observables are a canonical coordinate and momentum of a particle, the product of the uncertainties of the coordinate and momentum is given by

 $\Delta Q \cdot \Delta P > \frac{h}{2}$, where Q is coordinate and P is momentum.

This relation is called the Hersenberg's Principle of Uncertainty or Heisenberg's Uncertainty relation.

(1) The interpretation of this uncertainty relation is as follows, it is impossible to define at a given time both the position of the particle and its momentum to an arbitrary degree of accuracy When the lower limit imposed by above relation is reached, increasing the accuracy in the position (decreasing ΔR) implies that the accuracy in the momentum diminishes (increasing ΔP), and Vice Versa.

The limitation expressed by this uncertainty relation arises from the fact that the Planck constant h is not zero. It is the very small value of h (= 6.626 $\times 10^{-34}$ J.s) on the macroscopic scale that renders this limitation totally negligible in classical mechanics, but has considerable effect in microscopic world. For example, let us consider a dust particle with a diameter on the order of 1 µm and mass $M \approx 10^{-15} \text{ kg}$, having a speed $V = 10^{-3} \text{ m/s}$. Its momentum is then equal to

$$P = mv = 10^{-8} \text{ kg} \cdot \text{m/s}.$$

If its position is measured to within 0.01 µm, then the uncertainty op in the momentum must satisfy

$$\Delta P \ge \frac{f_{/2}}{\Delta R} = \frac{10^{-3k}}{10^{-8}} = 10^{-26} \text{ kg} \cdot \text{m/s}, \text{ i.e., } \frac{\Delta P}{P} \ge 10^{-8}.$$

Such a small uncertainty is totally negligible in the macroscopic World, as in practice, a momentum measurement device is incapable of attaining the required relative accuracy of 10^{-1} . In the macro world, we can regard $h \rightarrow 0$, so a particle can still have accurate position and momentum simultateously, i.e., a certain motion orbit.

Now let us consider the microscopic world, for example, the hydrogen atom. Assume the electron is moving along the n_1 orbit, and position uncertainty $\Delta X \sim a_1 = 0.53 \times 10^{-70} m$. The momentum uncertainty is given by

$$\Delta P \ge \frac{\hbar/2}{\Delta \chi} = \frac{6.626 \times 10^{-34}/2}{0.53 \times 10^{-10}} = 6.25 \times 10^{-24} \text{ kg} \cdot \text{m/s}.$$

However, the momentum of the electron itself is given by

$$P = me \mathcal{V}_{i} = me \sqrt{\frac{e^{2}}{4\pi\epsilon_{o}me a_{i}}} = e \sqrt{\frac{me}{4\pi\epsilon_{o}a_{i}}}$$

$$= 1.6 \times 10^{-19} \times \sqrt{\frac{9.1 \times 10^{-31}}{\sqrt{4\pi\times8.85 \times 10^{-12} \times 0.53 \times 10^{-10}}}$$

$$= 2.0 \times 10^{-24} \text{ kg} \cdot \text{m/s}.$$
Thus, $\frac{\Delta P}{P} \approx \frac{6.25 \times 10^{-24}}{2.0 \times 10^{-24}} \approx 3$, i.e., $\frac{\Delta P}{P} \geq 300\%$!
The relative uncertainty is so large that we cannot tell how much the electron momentum is !!!

(2) Derivation of Uncertainty Relation from commutation relation: First, We need a precise definition of the uncertainties of the coordinate Q and momentum P. The uncertainties ΔQ and Δp are defined as the root-mean-square deviation given by $\Delta Q = \sqrt{\langle (\hat{Q} - \langle \hat{Q} \rangle)^2 \rangle}$ where, $\langle \rangle$ "means taky mean, $\Delta P = \sqrt{\langle (\hat{P} - \langle \hat{P} \rangle)^2 \rangle}$ $P = \sqrt{\langle (\hat{P} - \langle \hat{P} \rangle)^2 \rangle}$

Now let us assume $|\psi\rangle$ is an arbitrary state but normalized, i.e., $\langle \psi|\psi\rangle = 1$. Operators correspondy to the coordinate and momentum are $\hat{\beta}$ and \hat{p} . Consider the ket vector

$$\begin{split} | \mathcal{P} \rangle &= (\hat{\alpha} + i\lambda \hat{p}) | \Psi \rangle, & \text{ where } \lambda \text{ is an arbitrary real number,} \\ \text{The conjugate of } | \mathcal{P} \rangle \text{ is } < \mathcal{P} | = | \mathcal{P}^* = <\Psi | (\hat{\alpha}^* - i\lambda^* \hat{p}^*) \\ \text{Since } \hat{\alpha} \text{ and } \hat{p} \text{ are Hermitian, i.e., } \hat{\alpha} = \hat{\alpha}^*, \ \hat{p} = \hat{p}^*, \ \text{we have} \\ <\mathcal{P} | = <\Psi | (\hat{\alpha} - i\lambda \hat{p}). \end{split}$$

For all λ , the square of the norm $\langle \varphi | \varphi \rangle$ is positive, i.e., $\langle \varphi | \varphi \rangle = \langle \psi | (\hat{\alpha} - i\lambda\hat{p}) (\hat{\alpha} + i\lambda\hat{p}) | \psi \rangle$ $= \langle \psi | \hat{\alpha}^2 + i\lambda\hat{\alpha}\hat{p} - i\lambda\hat{p}\hat{\alpha} + \lambda^2 \hat{p}^2 | \psi \rangle$ $= \langle \psi | \hat{\alpha}^2 | \psi \rangle + i\lambda \langle \psi | \hat{\alpha}\hat{p} - \hat{p}\hat{\alpha} | \psi \rangle + \lambda^2 \langle \psi | \hat{p}^2 | \psi \rangle$ $= \langle \hat{\alpha}^2 \rangle + i\lambda \langle \tau \hat{\alpha}, \hat{p} \rangle + \lambda^2 \langle \hat{p}^2 \rangle$ $\stackrel{\forall}{=} \langle \hat{\alpha}^2 \rangle - \hbar\lambda + \langle \hat{p}^2 \rangle \lambda^2 \geqslant 0.$ Recall the commutation velation $\tau \hat{\alpha}, \hat{p} \rangle = i\hbar.$ For polynomial expression $\langle \hat{\alpha}^2 \rangle - \hbar\lambda + \langle \hat{p}^2 \rangle \lambda^2 \geqslant 0,$

the discriminant of this expression must be negative or zero. [Discriminant for quadratic $ax^2+bx+c = 0$ is (b^2-4ac) .]

 $h^2 - 4 < \hat{\rho}^2 > < \hat{Q}^2 > \leq 0.$ Therefore, we have $\langle \hat{a}^2 \rangle \langle \hat{p}^2 \rangle \geq \frac{\hbar^2}{4}$ Now define $\hat{\alpha}' = \hat{\alpha} - \langle \hat{\alpha} \rangle, \quad \hat{P}' = \hat{P} - \langle \hat{P} \rangle.$ Then we find $\hat{\Theta}'$ and \hat{P}' have the same commutation relation as $\hat{\Theta}$ and \hat{P}' : $[\hat{\Theta}', \hat{P}'] = i\hbar$ $[\hat{\Theta}', \hat{P}'] = \hat{\Theta}'\hat{P}' - \hat{P}'\hat{\Theta}'$ $=(\hat{a}-\langle \hat{a} \rangle)(\hat{p}-\langle \hat{p} \rangle)-(\hat{p}-\langle \hat{p} \rangle)(\hat{a}-\langle \hat{a} \rangle)$ Recall < a > and < p > = $(\hat{a}\hat{p} - \hat{a} < \hat{p} > - < \hat{a} > \hat{p} + < \hat{a} > < \hat{p})$ are mean numbers, = $(\hat{a}\hat{p} - \hat{a} < \hat{p} > - < \hat{a} > \hat{p} + < \hat{a} > < \hat{p})$ Therefore, going through the same procedure as $\hat{\Theta} + i\lambda \hat{\rho}$, $|\mathcal{P}'\rangle = (\hat{\alpha}' + i\lambda \hat{P}')|\psi\rangle$, the square of the norm $\langle g' | \varphi' \rangle = \langle (\hat{\alpha}')^2 \rangle + i\lambda \langle c\hat{\alpha}', \hat{\rho}' \rangle + \lambda^2 \langle (\hat{\rho}')^2 \rangle$ $= \langle (\hat{a}')^2 \rangle - \hbar \lambda + \langle (\hat{P}')^2 \rangle \lambda^2 \geqslant 0$: $f^2 - 4 < (\hat{a}')^2 > < (\hat{P}')^2 > \leq 0$ Therefore, $\langle (\hat{\beta}')^2 \rangle \langle (\hat{\beta}')^2 \rangle \geq \frac{\hbar^2}{4}$. Accordy to the definition of uncertainty: $\Delta Q = \sqrt{\langle (\hat{Q} - \langle \hat{Q} \rangle)^2 \rangle} = \sqrt{\langle (\hat{Q}')^2 \rangle}$ $\Delta P = \sqrt{\langle (\hat{P} - \langle \hat{P} \rangle)^2 \rangle} = \sqrt{\langle (\hat{P}')^2 \rangle}$ Therefore, $\Delta Q \cdot \Delta P = \sqrt{\langle \hat{Q}' \hat{P} \rangle \langle (\hat{P}')^2 \rangle} \ge \sqrt{\frac{\hbar^2}{4}} = \frac{\hbar}{2}$.

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(3) The Heisenberg Uncentainty Principle can be generalized to the following: if two Observables are conjugates (like the coordinate and momentum are conjugates with each other), there exists an exact lower bound for the product of Uncertainties, which is equal to $\hbar/2$.

Further generalization of Uncertainty relation is that two arbitrary observables \hat{A} and \hat{B} have such limitation.

$$\Delta A \cdot \Delta B \ge \frac{1}{2} |\langle \Gamma \hat{A}, B \rangle|$$

In other words, the minimum bound of the product of Uncertainties is determined by the commutation velation between these two observables. If \hat{A} and \hat{B} are commutiz, i.e., $[\hat{A}, \hat{B}] = 0$, then $\Delta A \cdot xB \ge 0$. This means that it is possible to determine \hat{A} and \hat{B} precisely. But if $[\hat{A}, \hat{B}] \neq 0$, then there is a minimum limit on

the accuracy of simultaneously determining A and B.

4. The Intrinsic Nature of Uncertainty Principle is the Wave-Particle Duality of Matter (all material particles, photons and electromagnetic waves). It is a law of the nature, not due to measurement technologies or imperfection.

An Example for Principle of Uncertainty – Electron Diffraction

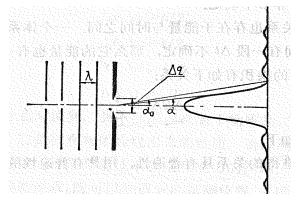


Diagram of electron diffraction (taken from Professor Chu's Atomic Physics)

Let the slit width be Δq . Since we cannot determine by which point of the slit the electron goes through, the position uncertainty of the electron is Δq .

Electron generates a diffraction pattern on the screen, so its motion direction can deviate from the original by angle of α . Thus, the momentum uncertainty $\Delta p = p \sin \alpha \approx p \alpha$. Let us use the first dark fringe position as a reference $\alpha_0 \approx \lambda/\Delta q$, then we have $\Delta p \approx p \alpha_0 = p \lambda/\Delta q$.

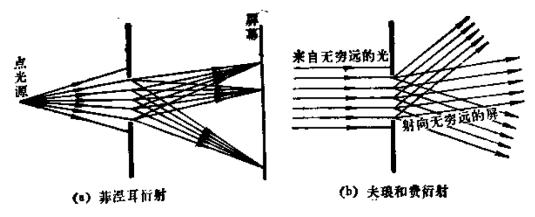
From de Broglie relationship, we know $\lambda = h/p$. By substituting this de Broglie wavelength into Δp equation, we get $\Delta p \approx p\lambda/\Delta q = h/\Delta q$. Therefore, we have

$$\Delta p \cdot \Delta q \sim h$$
.

This is larger than $\frac{\hbar}{2} = \frac{h}{4\pi}$, which is allowed by the uncertainty principle.

Note: 1) The unequal fringe heights in the double-slit experiment are due to diffraction modulating the interference fringes.

2) Diffraction's nature is multiple-beam interference, as shown in the diagram below.



Nature of diffraction is multiple-beam interference (from Professor Zhao's Optics)

An example for Uncertainty Principle: Electron Diffraction Heisenberg derived Ag Kala Kala Kala $\Delta g \cdot \Delta P \ge \frac{h}{z}, \quad h = \frac{h}{2\pi}$ resulted from the Wave-particle Electron Diffraction duality! $\frac{h}{2} = \frac{h}{4\pi}$ Assume the slit width be dg. Since we cannot determine by which point of the slit the electron goes through, the position uncertainty of the electron is 22. Electron generates a diffraction pattern on the Screen, so its movement direction can deviate the original direction angle or. => momentum uncertainty ap = psind = pd Use the 1st dark position as a reference $\alpha_0 \approx \lambda/\Delta g$ => Ap = Pa = P. 2/2g From de Broglie relationskip: $\lambda = h/p$ $\Rightarrow \Delta p = p \cdot \lambda / sg = \frac{h}{\Delta g}$ $\therefore \Delta P \cdot \Delta g \sim h > \frac{h}{2} = \frac{h}{4\pi}$ Note: the unequal fringes in the double-slit experiment pore due to the modulation of fringes by diffraction. Diffraction's nature is the multiple bean interference.

5. Application Examples of Uncertainty Principle (taken from Yang's Atomic Physics)

Example 1. The minimum mean kinetic energy of a bound particle

Assume a particle is confined within a range of r, so its position uncertainty is $\Delta x = r$. According to the principle of uncertainty, its momentum uncertainty has a minimum of

$$\Delta p_x = \frac{\hbar}{2\Delta x} = \frac{\hbar}{2r}$$
$$\Delta p_x = \sqrt{\langle (p_x - \overline{p}_x)^2 \rangle}$$

Recall the definition of Δp_x is

In a 3-dimentional space,

For a particle bound in certain space, its momentum in any direction should have a mean of zero, i.e., $\overline{p}_x = 0$. Therefore, we have

$$(\Delta p_x)^2 = < p_x^2 > .$$

$$< p_x^2 >= \frac{1}{3} < p^2 >$$

Therefore, in non-relativity range, the minimum mean kinetic energy is given by

$$E_k = \frac{\langle p^2 \rangle}{2m} = \frac{3\hbar^2}{8mr^2}.$$

It is obvious from this equation that the particle's kinetic energy cannot be zero! This conclusion does NOT depend on how the particle is confined. As long as a particle is in a potential well, its minimum kinetic energy cannot be zero. That is, the particle cannot fall to the bottom of the potential well. In fact, if the particle's kinetic energy is zero, then the principle of uncertainty tells us that its $\Delta x \rightarrow \infty$ -- how can the particle be confined?!

Example 2. Why can't the electrons inside an atom fall in the nucleus?

Bohr's theory cannot explain why the accelerated electron inside a hydrogen atom would not radiate electromagnetic waves (so losing energy) and would not fall in the nucleus. This question is answered by the principle of uncertainty in the following.

The electron inside an atom is bound spatially, so its kinetic energy can be given by the equation in Example 1. When the electron is getting closer to the nucleus, that is, r is getting smaller, Δx changes from the atomic dimension of 0.1 nm to the nuclear size of fm (10⁻¹⁵ m). According to the principle of uncertainty, the electron's momentum uncertainty increases,

$$E_k = \frac{\langle p^2 \rangle}{2m} = \frac{3\hbar^2}{8mr^2}$$

or using the equation in Example 1, its mean kinetic energy increases significantly. For example, when the electron motion range is decreased from 0.1 nm to 3 fm, its mean kinetic energy is increased from 1 eV to 0.1 GeV. Nowhere can the electron get such large energy; therefore, it is impossible for the electron to even get close to the nucleus, no mention fall inside the nucleus. A nucleus can have protons and neutrons, but not electrons, because the principle of uncertainty does not allow the confinement of electron within a nuclear size.

Example 3. Natural linewidth of a spectral line: Excited state lifetime vs. energy level width

From Bohr's theory, we know that when two energy levels corresponding to a spectral line have certain energy values, the transition between these two levels will give a spectral line whose width is infinitely narrow (just like a δ function), according to Bohr's frequency condition $hv = E_2 - E_1$. However, as long as an electron transitions from the upper level to the lower level, the upper energy level must have finite lifetime. Therefore, Δt is not infinite ∞ . According to the principle of uncertainty, $\Delta E \neq 0$, i.e., the upper energy level must have certain width ΔE . Consequently, the spectral line cannot be a δ function, but must have certain linewidth. This is called the natural linewidth of the spectral line. For example, if an excited state has a lifetime of $\Delta t = 10^{-8}$ s, then

$$\Delta E \ge \frac{\hbar}{2\Delta t} = 3.3 \times 10^{-8} eV.$$

The spectral linewidth can be derived as $\Delta v_n = \frac{\Delta E}{h} \ge \frac{\hbar}{2h\Delta t} = \frac{1}{2 \cdot 2\pi \cdot \Delta t}$.

Natural linewidth calculation: In reality (taken from Professors Chu and Papen lidar book chapter), the natural linewidth is given by, for Na D₂ line (3 ${}^{2}P_{3/2} \rightarrow 3 {}^{2}S_{1/2}$), $\tau = 16.23$ ns

$$\Delta v_n = \frac{1}{2\pi \cdot \tau} = \frac{1}{2\pi \times 16.23ns} = 9.8 MHz$$

For a metastable state, the lifetime is much longer, e.g., on the order of μ s, ms, or s. Then its natural linewidth can be significantly narrower than MHz.

$$\Delta E \cdot \Delta t = \frac{\hbar}{2}$$

$$E = mc^{2} = \int P^{2}c^{2} + m_{0}^{2}c^{4} \qquad - \text{Relativity Range}$$

$$\partial E = \frac{1}{2} \not(P^{2}c^{2} + m_{0}^{2}c^{4})^{\frac{1}{2}} \quad 2c^{2}P \, dP$$

$$= \frac{c^{2}P}{E} dP = \frac{P}{m} \frac{dP}{m} = \mathcal{V} dP$$
Non-relativity vange: $E_{K} = \frac{1}{2}m\mathcal{V}^{2}, P = m\mathcal{V}^{-1}$

$$\therefore E_{K} = \frac{P^{2}}{2m} \Delta P = \mathcal{V} \cdot \Delta P.$$

$$\therefore \Delta E \cdot \Delta t = \Delta t \mathcal{V} \cdot \Delta P = \Delta g \cdot \Delta P \ge \frac{\hbar}{2}$$
An example, Excited state lifetime vs. energy level width metastable
$$\Delta U = \frac{\Delta E}{h} \ge \frac{\hbar}{2\pi \Delta t}$$

$$\Delta U = \frac{\Delta E}{h} \ge \frac{\hbar}{2\pi \Delta t} \cdot h$$

$$= \frac{1}{4\pi \cdot \Delta t}$$
To reality,
$$\Delta U_{h} = \frac{1}{2\pi \cdot \Delta t}$$
For example, $Na({}^{2}P_{3}h_{2}) = 16.33ns \Longrightarrow \Delta y_{h} = 9.8 \text{ MM}_{H}$

$$Fe$$

§3.5. Equation of Motion – Schrödinger Equation

1. Introduction to the Principle of Motion – the Schrödinger equation describes the time evolution of the state of a quantum system.

- 1) Schrödinger equation is a fundamental equation in QM. In principle, it cannot be derived or proven from more fundamental principles, but it can only be verified by experiments.
- 2) Hamiltonian operator
- 3) Schrödinger equation in the abstract state space as well as in any representation.
- 4) There is no indeterminacy in the time evolution of a quantum system. Indeterminacy appears only when a physical quantity is measured, the state vector undergoes an unpredictable modification, i.e., the reduction or collapse of the state to an eigenstate. However, between two measurements, the state vector evolves in a perfectly deterministic way, in accordance with the Schrödinger equation.
- 5) The Schrödinger equation is linear, so its solution is linearly superposable.
- 6) Time evolution does not modify the global probability of finding the particle in all space, which always remains 1.

2. Stationary-State Schrödinger Equation – Energy Eigenvalue Equation

3. Application Examples of Schrödinger Equation in Solving Atomic Particle Problems

Three examples are presented to demonstrate

- 1) How to solve stationary-state Schrodinger equations
- 2) How Schrodinger equation naturally leads to quantized energy, etc.
- 3) Real applications of quantum mechanics in modern physics and technologies

Example 1. 1-D Infinite Potential Well

- Points: Boundary conditions naturally lead to the quantized energy levels and the minimum energy isn't zero; wave function represents a standing wave; the distribution of probability of particle isn't homogeneous
- Applications: Hydrogen atom's electron is similar to be in quantum well; quantum well in laser, photo detector, etc.

Example 2. Potential Barrier Penetration (Finite Potential Well) → quantum tunnel effect

- Points: Particles have probability appearing outside classical physics barriers
- Applications: emission of α particle by radio-active nucleus, tunnel diode, electron tunneling microscope

Example 3. 1-D Harmonic Oscillator

- Points: Quantized and non-zero energy of oscillator
- Applications: Quantized radiation field

\$3.5 Equation of Motion - Schrödinger Equation

14. QM principle of motion describes the time evolution of the state of a system, which is governed by the Schrödinger Equation. $i\hbar \frac{d}{dt} | \Psi(t) \rangle = \hat{H}(t) | \Psi(t) \rangle$

where $|\Psi(t)\rangle$ is the state vector, $\hat{H}(t)$ is the observable associated with the total energy of the system (called the Hamiltonian operator of the system), \hat{T} is planck constant divided by 2TT, and $\frac{d}{dt}$ is to take derivative of time. (1) The Schrödinger equation is a fundamental equation for the Quantum Mechanics, just like the Newtonian equation in the Classical mechanics. In principle, it cannot be derived or proven from more fundamental principles, but it can only be verified by experiments. The Schrödinger equation describes the change of state with time, the time evoluation of motion, or the change of motion. (2) The Hamiltonian operator \hat{H} comes from the classical Hamiltonian: $H(t) = \frac{\hat{P}(t)}{2m} + V(t)$,

Where $\vec{P}(t)$ is the momentum, m is the mass, and V(t) is the potential energy of the system. \vec{P}^{z} represents the Kinetic energy of the System. In OM, the Hamiltonian Operator is given by $\vec{H}(t) = \frac{\vec{P}(t)}{2m} + V(t)$. (3) Note that the Schrödinger equation is expressed in the abstract state space with an abstract operator \vec{H} , abstract State vector $|V(t)\rangle$, and a full derivative \vec{d} . When express the Schrödinger Equation in any representation, it will become a partial derivative $\frac{\partial}{\partial t}$, and $\hat{H}(t)$ and $|\Psi(t)\rangle$ will also be projected to this representation. The equation may not be such simple format, depending on what representation is chosen. (4) The Schrödinger equation is of first order int. From this it follows that, given the initial state $|\Psi(t_0)\rangle$, the state $|\Psi(t)\rangle$ at any subsequent time t is determined. There is no

(V(t)) at any subsequent time t is determined. There is no in determinacy in the time evolution of a guantum System. Indeterminacy appears only when a physical guantity is measured, the state vector then undergoing an unpredictable modification (i.e., the reduction of the state to an eigenstate). However, between two measurements, the state vector evolves in a perfectly deterministic way, in accordance with the Schrödinger equation.

(5) The Schrödinger Equation is linear and homogeneous. It follows that its solution is linearly superposable.

Let $|\Psi_{i}(t)\rangle$ and $|\Psi_{i}(t)\rangle$ be two solutions of the Schrödiger equation: if $f_{t}|\Psi_{i}\rangle = \hat{H}|\Psi_{i}\rangle$

 $i\hbar \frac{d}{dt} | \psi_2 \rangle = \hat{H} | \psi_2 \rangle.$

Then the linear superposition of $|4,(t)\rangle$ and $|4_2(t)\rangle$. $|4/(t)\rangle = \lambda_1 |4_1(t)\rangle + \lambda_2 |4_2(t)\rangle$ (λ_1 and λ_2 are two complex constants) is also a solution to the Schrödinger Equation: $i\hbar \frac{d}{dt} |4(t)\rangle = \hat{H} |4(t)\rangle$ The statement can be verified as below:

$$i\hbar \frac{d}{dt} | 4 \rangle = \lambda, i\hbar \frac{d}{dt} | 4 \rangle + \lambda_2 i\hbar \frac{d}{dt} | 4_2 \rangle$$

$$= \lambda, \hat{H} | 4 \rangle + \lambda_2 \hat{H} | 4_2 \rangle$$

$$= \hat{H} (\lambda, | 4 \rangle + \lambda_2 | 4_2 \rangle)$$

$$= \hat{H} | 4 \rangle$$

Thus, the superposition $|\psi\rangle$ is a solution to the Schrödinger equation. (6) The norm of the state vector $|\psi(t)\rangle$ is defined as $\sqrt{\langle \psi(t)|\psi(t)\rangle}$. Since the Hamiltonian operator $\hat{H}(t)$ is a Hermitian, i.e., $\hat{H}(t) = \hat{H}^*(t)$, the square of the norm of the state vector, $\langle \psi(t)|\psi(t)\rangle$, does not depend on t as we can see below:

 $\frac{d}{dt} \langle \Psi(t) | \Psi(t) \rangle = \left[\frac{d}{dt} \langle \Psi(t) | \right] | \Psi(t) \rangle + \langle \Psi(t) | \left[\frac{d}{dt} | \Psi(t) \rangle \right].$ From the Schrödinger equation, we have $\frac{d}{dt} | \Psi(t) \rangle = \frac{1}{i\hbar} \hat{H} | \Psi(t) \rangle,$ and $\frac{d}{dt} \langle \Psi(t) | = -\frac{1}{i\hbar} \langle \Psi(t) | \hat{H}^* = -\frac{1}{i\hbar} \langle \Psi(t) | \hat{H}.$ Substituting these two items into above derivative equation, we have

$$\frac{d}{dt} < \psi(t) | \psi(t) \rangle = -\frac{1}{th} < \psi(t) | \hat{H} | \psi(t) \rangle + \frac{1}{th} < \psi(t) | \hat{H} | \psi(t) \rangle$$
$$= 0$$

Therefore, $\langle \Psi(t) | \Psi(t) \rangle = \langle \Psi(t_0) | \Psi(t_0) \rangle = 1$. The last equality comes if the state vector $| \Psi_0(t) \rangle$ is normalized at time to.

The property of the norm conservation is very useful in guantum mechanics. For example, it becomes indispensable when we interpret the square of the modulus $|\langle \psi(\vec{r},t) \rangle^2 = |\langle \vec{r} | \psi \rangle|^2$ of the wave function of a spinless particle as being the position probability density. Time evolution does not modify the global probability of finding the particle in all space, which alway remains equal to 1. Therefore, $\langle \psi(t) | \psi(t) \rangle = (d^3 r | \psi(\vec{r},t) |^2 = 1.$

\$3.6 Abstract (Space, State Vectors, Linear Operators Vs. Representations

Abstract state Space, state vectors, linear operators Representations, choosing an orthonormal basis, then projecting the state vector and linear operator onto the basis so that they are represented by numbers or functions. - Convenient for calculation.

1) State Vector

2) linear operator 3) Mean Value Cakulation

$$= \text{Commutation Relation} \\ \hline Commutation Relation : $[X, P_X] = i\hbar \\ F\hat{Y}_i, \hat{P}_i] = i\hbar \hat{J}_i \quad (i, j = X, Y, z)$$$

(Dirac) \$3.6 Abstract Space Versus Representations

15. Representations:

Above descriptions of QM concepts and principles have been using the abstract state vectors and the abstract linear operators in the abstract state space. They are perfect for expressing the exact meaning of QM principle of superposition of states, principle of uncertainty, principle of motion, and principle of measurements. They also give clear presentation of some derivations like projector operator, mean value, etc.

2)

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However, when facing practical problems, we need to project these abstract vectors and operators to a concrete representation in order to do Some actual computation. Choosing a representation means choosing an orthonormal basis, either discrete or continuous, in the state space. We then project the state vector and linear Operator onto the basis so that the state vector and linear Operator are represented by numbers or functions in the actual representation. The choice of a representation is, in theory, arbitrary. In practice, it obviously depends on the particular problem being studied: in each case, one chooses the representation which leads to the simplest Calculation.

Two important examples of representations and observables are the coordinate (position) $\{|\vec{r}\rangle\}$ and momentum $\{|\vec{P}\rangle\}$ representations and observables (\vec{r} and \vec{P}). We will use these two as examples to show how to project state and operator to a representation, how to do actual calculation, and how to change from representation to another.

(1) State vector versus Wave function in
$$\{I\vec{P}\}\)$$
 and $\{I\vec{P}\}\)$ representation
Recall the principle of superposition of states, for the complete
orthonormal states $\{I\vec{P}\}\)$, we have $\langle\vec{r}'|\vec{r}\rangle = \delta(\vec{r}-\vec{r}')$.
Similarly, for orthonormal $\{I\vec{P}\}\)$, we have $\langle\vec{P}'|\vec{P}\rangle = \delta(\vec{P}-\vec{P}')$.
Y Any arbitrary state vector $|W\rangle$ can be expanded in terms of
 $\{I\vec{r}\}\)$ as: $|W\rangle = \int C(\vec{r}')|\vec{r}'\rangle d^3r'$, where $C(\vec{r}')$ is the
probability amplitude of the system being in the state of $|\vec{r}'\rangle$,
i.e., $|C(\vec{r}')|^2$ is the probability of findy the system at position
 \vec{r}' . As explained earlier,
 $\langle\vec{r}'|W\rangle = \int C(\vec{r}') \langle\vec{r}|\vec{r}'\rangle d^3r' = \int C(\vec{r}')\delta(\vec{r}-\vec{r}')d^3r' = C(\vec{r}')$
Thus, we repeat the expression for $C(\vec{r}')$ [i.e., $C(\omega)$].
We can write $C(\vec{r}') = \langle\vec{r}|W\rangle \equiv \Psi(\vec{r}')$,
which is called the Wave function in the $\{I\vec{P}\}\)$ representation.
* Similarly, $|W\rangle$ can be projected onto $\{I\vec{P}>\)$ representation.
* Similarly, $|W\rangle$ can be projected onto $\{I\vec{P}>\)$ representation.
* Wave function $\Psi(\vec{r}) = \langle\vec{r}|W\rangle = \psi(\vec{P})$,
which is the wave function in the $\{I\vec{P}>\)$ representation.
* Wave function $\Psi(\vec{r}) = \langle\vec{r}|W\rangle$ has the following mean \vec{r} . It is
the probability amplitude of finding a particle at position \vec{r} .
In other words, $|\Psi(\vec{P})|^2 = |\langle\vec{r}|W\rangle|^2$ is the probability of the
particle appearing at position \vec{r} . Similarly, $|\Psi(\vec{P})|^2 = |\langle\vec{P}|\Psi|^2$
is the probability of finding the particle with waventum \vec{P} .

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* Wave function
$$\psi(\vec{r})$$
 is the projection of the state vector $|\Psi\rangle$
onto the $\{1\vec{r}\}\)$ representation, analogy to a vector \vec{r} being
projected onto $a(X, Y, z)$ coordinate space:
 $\vec{r} = X\vec{i} + \forall \vec{j} + z\vec{k}$ $(\vec{i}, \vec{j}, \vec{k} \text{ are unit}$
 $vectors along$
 $\vec{r} \Rightarrow (X, \forall, 2)$. $X, \forall, z \text{ axes}$)
Analogy: $|\Psi\rangle = \int \Psi(\vec{r}) |\vec{r}\rangle d^{3}r$
 $i |\Psi\rangle \Rightarrow \{\Psi(\vec{r})\}$
* In 1-dimension case, $\vec{r} \rightarrow X$, so $|\Psi\rangle \Rightarrow \{\Psi(x)\}$.
The conjugate of $|\Psi\rangle$, i.e., $\langle\Psi| \Rightarrow \{\Psi^{*}(x)\}$. These are
the wave functions used in HW #2, problem #2.
* Note $|\langle\vec{r}|\Psi\rangle|^{2} \equiv \langle\vec{r}|\Psi\rangle^{*} \langle\vec{r}|\Psi\rangle$
 $|\Psi(\vec{r})|^{2} \equiv \langle\Psi(\vec{r})\Psi(\vec{r})$.

* In principle, different representations are equivalent, since they represent the same state vector 14> and same operator A.

(2) Operator in {IT>} and fIP>} representations In {IT>} representation, the coordinate T and momentum P operators are $X \to X$, $\frac{\Lambda}{\Gamma} \to \overline{\Gamma}$ $\hat{p}_{x} \rightarrow -i\hbar\hat{g}_{x}$, $\hat{p} \rightarrow -i\hbar\vec{r}$, $\vec{p} = \vec{e}_{x}\hat{g}_{x} + \vec{e}_{y}\hat{g}_{y} + \vec{e}_{z}\hat{g}_{z}$ In {IP>} representation, f and p operators are $\vec{R} \rightarrow \vec{R}, \vec{P} \rightarrow \vec{P}$ ネーボー、デーボー、デー・気気+気が、 Because V(P) and V(P) are the Fourier transform of each other, we have $\hat{p} = -i\hbar \vec{p}$ in $\{|\vec{r}\rangle\}$ representation, while $\hat{\gamma} = i\hbar \vec{\nabla}_p$ in { $|\vec{P}\rangle$ } representation. * {IT>> is the most common and widely used representation, also Called the Schrödinger representation, in Which Schrödinger developed the Schrödinger equation. Let us write down a few other operators in the fir>3 representation. D Kinetic energy operator, the classical kinetic energy is given by $E_{k} = \frac{\overline{P}^{2}}{zm}$. To guartize it to derive its operator, replace p with its correspondy operator p: $\vec{E}_{k} = \frac{\vec{P}^{2}}{2m} = -\frac{\hbar^{2}}{2m} \left(\frac{\partial^{2}}{\partial \chi^{2}} + \frac{\partial^{2}}{\partial y^{2}} + \frac{\partial^{2}}{\partial z^{2}} \right) = -\frac{\hbar^{2}}{2m} \nabla^{2}.$ (2) Hamiltonian operator $\hat{H}(t)$: the classical $H = \frac{\vec{p}^2}{2m} + V$ Then in QM, $\hat{H}(t) = \frac{\vec{p}^2}{2m} + \hat{V}$. If $\hat{V} = V(\vec{r})$, then $\hat{V} = V(\vec{r})$ Since $\vec{r} = \vec{r}$ in $\{I\vec{r}\}\}$, $\hat{H}(t) = -\frac{\hbar^2}{2m} \left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2}\right) + V(\vec{r})$.

(3) Angular momentum operator: the classical angular momentum
is
$$\vec{l} = \vec{r} \times \vec{p}$$
, where \vec{r} is position vector, \vec{p} is momentum vector.
Corresponding QM operator is:
 $\hat{\vec{l}} = \hat{\vec{r}} \times \hat{\vec{p}} = -i\hbar \vec{r} \times \vec{V}$.
Angular momentum operator components are
 $\left(\hat{l}_x = -i\hbar (y \frac{\partial}{\partial z} - 2 \frac{\partial}{\partial y}) = i\hbar (sing \frac{\partial}{\partial \theta} + cot\theta \cos g \frac{\partial}{\partial g})\right)$
 $\hat{l}_y = -i\hbar (2\frac{\partial}{\partial x} - x \frac{\partial}{\partial z}) = i\hbar (-\cos g \frac{\partial}{\partial \theta} + \cot \theta \sin g \frac{\partial}{\partial g})$
 $\hat{l}_z = -i\hbar (x \frac{\partial}{\partial y} - y \frac{\partial}{\partial x}) = -i\hbar \frac{\partial}{\partial q}$
The last equality is in spherical coordinates:
 $r = \sqrt{x^2 + y^2 + z^2}, \ \theta = \arctan(\sqrt{x^2 + y^2}/2), \ \theta = \arctan(y/x) \frac{g}{g}$.
(3) The square of angular momentum:
 $\vec{l}^2 = \hat{l}_x^2 + \hat{l}_y^2 + \hat{l}_z^2$
 $= -\hbar^2 [\frac{1}{\sin \theta} \cdot \frac{\partial}{\partial \theta} (\sin \theta \frac{\partial}{\partial \theta}) + \frac{1}{\sin^2 \theta} \frac{\partial^2}{\partial g^2}]$.

.

(3) Mean values in the fIP>3 and fIP>3 representations. In the abstract state space, as we proved above, the mean value of an observable A in the state 14> of a system is given by $\overline{A} \equiv \langle \widehat{A} \rangle = \langle \Psi | \widehat{A} | \Psi \rangle.$ Now project state vector and operator to {17>} representation, recall $\int |\vec{r}\rangle < \vec{r} | d^3r = 1$, we have $\overline{A} = \langle \psi | \hat{A} | \psi \rangle$ $= \langle \psi | \left(\int |\vec{r} \rangle \langle \vec{r} | d^3 r \right) \hat{A} | \psi \rangle$ $= \int d^3r < \psi |\vec{r} > < \vec{r} |\hat{A}| \psi > .$ Recall $\langle \Psi|\vec{r}\rangle = \Psi(\vec{r})$. If \hat{A} is a function of \vec{r} , i.e., $\hat{A} = A(\vec{r})$, then since $|\vec{r}\rangle$ is an eigenstate of $\hat{\vec{r}}$, it is also an eigenstate of $A(\vec{r})$, i.e., $A(\vec{r})|\vec{r}\rangle = A(\vec{r})|\vec{r}\rangle$. Thus, $\langle \vec{F} | \hat{A} | \psi \rangle = \langle \vec{F} | A c \hat{F} \rangle | \psi \rangle = A (\vec{F}) \langle \vec{F} | \psi \rangle = A (\vec{F}) \langle \vec{F} | \psi \rangle = A (\vec{F}) \langle \vec{F} | \psi \rangle$ $A = \int d^3r \, \psi^*(\vec{r}) \, A(\vec{r}) \, \psi(\vec{r}) \, .$ Example: In HW#2, Problem #7(1),1-D case, X2 is a function of \hat{X} , i.e., $\hat{X}^2 = (\hat{X})^2$ $\therefore \hat{X}|X > = X|X > \therefore \hat{X}^2|X > = X^2|X >$ $\int \sqrt{X^2} = \int dx \ (\psi(x) \ X^2 \ \psi^*(x)).$ Furthermore, since X² is just a polynomial term of X, it can exchange order with (t*cx) or ((cx), so $\overline{\chi^2} = \left(\chi^2 |\psi(x)|^2 dx\right).$

For the momentum operator $\hat{\vec{p}} = -i\hbar \vec{p}$ in $S_1\vec{r}>g$ representation, which is NOT a simple function of \vec{r} , but involves derivatives.

$$\overline{P} = \langle \widehat{P} \rangle = \langle \psi | \widehat{P} | \psi \rangle$$

$$= \langle \psi | (\int | \overrightarrow{P} \rangle \langle \overrightarrow{r} | d^{3}r) (-i\hbar \overrightarrow{P}) | \psi \rangle$$

$$= \int d^{3}r \langle \psi | \overrightarrow{P} \rangle (-i\hbar) \langle \overrightarrow{r} | \overrightarrow{P} | \psi \rangle.$$

It can be proven that $\langle \vec{r} | \vec{\beta} | \psi \rangle = \vec{\rho} \langle \vec{r} | \psi \rangle = \vec{\rho} \langle \vec{r} \rangle$ $\Rightarrow \vec{\rho} = \int d^3r \ \psi(\vec{r}) \ (-i\hbar \vec{\sigma}) \ \psi(\vec{r}).$

If an operator is a function of \vec{P} , e.g., $\vec{B} = B(\vec{P})$, then $\hat{B}(\vec{r}) = B(-i\hbar\vec{P})(\vec{r})$

$$\vec{B} = \int d^3r \, (\psi \vec{c} \vec{r}) B(-i\hbar \vec{r}) \, \psi(\vec{r})$$

Example, In HW #2, Problem #7, 1-D case, $\hat{P}_{x} = -i\hbar\frac{\partial}{\partial x}$, $80(\hat{P}_{x} - P_{o})^{2} = (-i\hbar\frac{\partial}{\partial x} - P_{o})^{2} = -\hbar^{2}\frac{\partial^{2}}{\partial x^{2}} + 2i\hbar P_{o}\frac{\partial}{\partial x} + P_{o}^{2}$.

 $(\hat{P}_{x} - P_{0})^{2} = \int dx \, (\ell^{*}cx) \left[-\hbar^{2} \frac{\partial^{2}}{\partial x^{2}} + 2i\hbar P_{0} \frac{\partial}{\partial x} + P_{0}^{2} \right] \, (\varphi cx) .$ Note, since $\frac{\partial^{2}}{\partial x^{2}}$, $\frac{\partial}{\partial x}$ are derivatives, not a polynomial of x, they <u>CANNOT</u> switch order with $\ell^{*}cx$ or $\ell(cx) !!!$

As $\psi(x) = A e^{-\alpha x^{2}/2} e^{+i\beta x/\hbar}$, $\psi^{*}(x) = A e^{-\alpha x^{2}/2} e^{-i\beta x/\hbar}$,

You cannot cancel out the imaginary part <u>BEFORE</u> taky derivatives, (Of course, after taky derivatives, you only have x polynomial terms left, then you can swith these x terms with $\Psi^{*}(x)$ or $\Psi(x)$, and cancel out $e^{\pm i P_{0} \times / \hbar}$ part !)

If you follow this procedure, problem #2 (4) should reach a correct answer $\Delta p = \hbar \sqrt{\frac{a}{2}}$.

The same state and same operator can also be projected to different vepresentations, e.g., SIP>3 presentation, and they are equivalent. The momentum operator $\vec{p} = \vec{p}$ in the { $|\vec{p}>$ } representation, · アミくデ> = < 41 デーレ> $= \langle \Psi | (\int d^3 p | \vec{P} \rangle \langle \vec{P} |) \vec{P} | \Psi \rangle$ $= \int d^3 p < \langle \psi | \vec{P} \rangle < \vec{P} | \vec{P} | \psi \rangle$ = [d3p < 4] P> P<P14> To distinguish <PIU> from <PIU>, we write <PIU>=CCPJ. $\therefore \vec{P} = \int d^3 p \cdot c^*(p) \vec{P} \cdot c(p)$ Similarly, if $\hat{B} = B(\hat{P})$, then $\overline{B} = \int d^{3}P \cdot C^{*}(P) B(\overrightarrow{P}) C(P).$ Example: In HOU #2, Problem #7 (2), C(PR) is the Fourier transform of $\psi(x)$, and $\hat{R} = P_x$, $(\hat{P}_x - P_o)^2 = (P_x - P_o)^2$. $(\hat{R} - P_0)^2 = \int dP_x (\hat{C}P_x) (P_x - P_0)^2 C(\hat{R}) = \hbar \sqrt{\frac{\alpha}{2}}$ It gives the same result as problem # 2(4). This indicates that the {IP>} representation is equivalent to the {IP>}. With the same state 14> and same operator \hat{p} , you Will get the same results (the root-mean-square) in the {IP>} and {IP>} representations. In other Words, the results only depend on the state vector 14> and the Observable operator but independent of representations. But choosy different representations makes the calculation easier or complicated just like poperator has more straightforward calculation in SIP>3.

(4) Eigenvalue Equation in
$$\{1\vec{r}\}\$$
 representation.
In the abstract state space, an eigenvalue equation is
 $\hat{A} \mid \psi \rangle = a \mid \psi \rangle$
Where $\hat{A} - observable operator, \quad |\psi \rangle - state vector,$
 $a - a complex$ constant (veal number in reality)
It can be proven that in the $\{1\vec{r}\}\$ representation,
the eigen value equation becomes
 $\hat{A} \mid \psi \mid \vec{r} \rangle = a \mid \psi \mid \vec{r} \rangle$
Where \hat{A} is the representation of \hat{A} operator $\hat{A} \mid \vec{r}\mid \vec{r} \rangle$.
 $\hat{A} \mid \vec{r}\mid \vec{r} \rangle$ and $\hat{A} \mid \vec{r}\mid \vec{r} \rangle$
where \hat{A} is the representation of \hat{A} operator $\hat{A} \mid \vec{r}\mid \vec{r} \rangle$.
 $\hat{A} \mid \vec{r}\mid \vec{r} \rangle$ representation of \hat{A} operator $\hat{A} \mid \vec{r}\mid \vec{r} \rangle$.
 $\hat{A} \mid \vec{r}\mid \vec{r} \rangle$ representation.
Example: $v = 0 \mid \vec{r} \mid \vec{r} \rangle$ representation.
Example: $v = 0 \mid \vec{r} \mid \vec{r} \rangle$ between $x = 0 = 1 \times a$. The normalized
 $\vec{r} \mid \vec{r} \rangle$ between $x = 0 = 1 \times a$. The normalized
 $\vec{r} \mid \vec{r} \rangle$ between $x = 0 = 1 \times a$. The normalized
 $\vec{r} \mid \vec{r} \rangle$ between $x = 0 = 1 \times a$. The normalized
 $\vec{r} \mid \vec{r} \rangle$ between $x = 0 = 1 \times a$. The normalized
 $\vec{r} \mid \vec{r} \rangle$ is the mean of 0 . elsewhere
the particle's momentum and kinetic energy
(2) Verify whether this wave function is an eigenfunction of
momentum, whether an eigen wave function of kinetic energy.
Solution: In $\{1x\}$ representation, $\hat{\vec{r}} = -i\hbar \frac{\partial}{\partial x}$, $\frac{\hat{\vec{r}}\mid^2}{\hat{r}\mid^2} = -\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x}$.
 $\vec{r} = -i\hbar \frac{Q}{a} \cdot \frac{m\pi}{a} \int_{0}^{a} \sin(\frac{m\pi}{a}) c_{q}(\frac{M\pi}{a}) dx$
 $= -i\hbar \frac{Q}{a} \cdot \frac{m\pi}{a} \int_{0}^{a} \sin(\frac{m\pi}{a}) c_{q}(\frac{M\pi}{a}) dx$

Kinetic energy mean:

Verifying eigen wave function: "Apply the operator to the wave function to see whether you can obtain a number times the same wave function " $(1) \hat{p} \Psi_n(x) = -i\hbar \frac{\partial}{\partial x} \Psi_n(x) = -i\hbar \sqrt{2} \frac{\partial}{\partial x} \sin\left(\frac{m\pi x}{a}\right) = i\sqrt{2} \frac{\hbar m}{a} \cos\left(\frac{m\pi x}{a}\right)$ - Unicx) is NOT p's eigen function. $(z) \frac{\hat{\rho}^2}{2m} \left(\ell_n(x) = -\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} \ell_n(x) = -\frac{\hbar^2}{2m} \sqrt{\frac{2}{\alpha}} \frac{\partial^2}{\partial x^2} \operatorname{Sin}(\frac{n\pi x}{\alpha}) \right)$ $=\frac{\hbar^2}{2m}\sqrt{\frac{2}{\alpha}\left(\frac{n\pi}{\alpha}\right)^2}\operatorname{Sin}\left(\frac{n\pi x}{\alpha}\right)=\frac{\hbar^2\pi^2n^2}{2m\alpha^2}\psi_n(x)$: (In (x) is the eigen function of kinetic energy. The eigenvalue is $\frac{\hbar^2 \pi^2 n^2}{2m n^2}$ the same as the mean we derived above - of course, the mean of the Kinetic energy in its own eigenstate is equal to the eigenvalue, Thus, we solved the HW#3, problem #3. You may practice it yourself. Then apply similar technique to HW#3, Problem#2.

(5) Schrödinger Equation in
$$\{1\vec{P}\}\)$$
 and $\{1\vec{P}\}\)$ representations.
In the abstract state space, the Schrödinger equation is
it $\frac{d}{dt} | \Psi(t) \rangle = \hat{H}(t) | \Psi(t) \rangle$.
For a (spinless) particle in a scalar potential $V(\vec{r})$, the
operator $\hat{H} = \frac{\hat{P}^2}{2m} + V(\hat{r})$.
It can be proven that in the $\{1\vec{r}\}\)$ representation, the
Schrödinger equation becomes:
 $i\hbar \frac{\partial}{\partial t} \Psi(\vec{r},t) = [-\frac{\hbar^2}{2m} \nabla^2 + V(\hat{r})] \Psi(\vec{r},t)$.
In the $\{1\vec{P}\}\)$ representation, the Schrödinger equation is
 $i\hbar \frac{\partial}{\partial t} \Psi(\vec{r},t) = \frac{\vec{P}^2}{2m} \overline{\Psi}(\vec{P},t) + (2\pi\hbar)^{\frac{3}{2}} \int d^3p' V(\vec{P},\vec{r})$.
Here, $\Psi(\vec{r},t) \equiv \langle \vec{r} | \Psi \rangle$.
 $\overline{V}(\vec{P}) = (2\pi\hbar)^{\frac{3}{2}} \int d^3r \ e^{\frac{\pi}{2}\vec{P}\cdot\vec{r}} V(\vec{r})$.

\$3.7 Solutions to Eigenvalue Equation and Schrödiger Equation

16. Solution to Eigenvalue Equation and Schrödinger Equation. Schrödinger equation: $i\hbar \frac{2}{3t} \psi(\vec{r},t) = \left[-\frac{\hbar^2}{2m} \nabla^2 + V(\vec{r})\right] \psi(\vec{r},t)$. If $V(\vec{r})$ is not explicitly dependent on time t, then we have $\psi(\vec{r},t) = \psi(\vec{r}) T(t)$.

Substituty this into the schrödinger equation:

$$\frac{i\hbar}{T} \frac{dT}{dt} = \frac{1}{\psi(\vec{r})} \left[-\frac{\hbar^2}{2m} \nabla^2 + V(\vec{r}) \right] \psi(\vec{r}) = E$$

$$\frac{i\hbar}{T} \frac{dT}{dt} = E \implies T = T_0 e^{-iEt/\hbar} \qquad T$$

$$\int \left[-\frac{\hbar^2}{2m} \nabla^2 + V(\vec{r}) \right] \psi(\vec{r}) = E \psi(\vec{r})$$

$$\frac{1}{2m} \left[-\frac{\hbar^2}{2m} \nabla^2 + V(\vec{r}) \right] \psi(\vec{r}) = E \psi(\vec{r})$$

Probability density = $|\psi(\vec{r},t)|^2 = |\psi(\vec{r})|^2$ is independent of t. i.e., the probability of the particle appearing at position \vec{r} does not change with time! Equation $\int -\frac{\hbar^2}{2m} \nabla^2 + \vec{V}(\vec{r}) = E \psi(\vec{r})$

(1) I-Dimension Infinite potential Well
The stationary-state Schrödinger equation

$$I = \frac{\pi^2}{2m} \frac{3}{2N^2+} V (Y) = E (Y)$$
.
Since a particle cannot be in an infinite.
 $V(x) = \begin{cases} w, x < 0 \\ 0, 0, x < a \end{cases}$
potential, $(\psi = 0 \text{ in Pegions II and II. (Physics)} \\ w, x > a \end{cases}$
In region I, $V = 0$, the equation is Simplified to
 $-\frac{\pi^2}{2m} \frac{3^2}{2N^2} (Y(x)) = E (Y(x))$
Let $K = \sqrt{\frac{2\pi}{\pi^2}}$, then $\frac{d^2 \psi}{dx} + K^2 \psi = 0$
The general solution to this equation is
 $\psi = A \sin (kx + \delta)$,
where A and δ are constants to be determined from boundary
conditions and intermedization requirements.
Considering from physics aspects, since the particle cannot be
In the $V = \omega$ region, i.e., the probability to be in regions
II and II is zero. Therefore, $\psi(x=0) = 0$, $\psi(x=a) = 0$.
At $x = 0$, $0 = A \sin \delta$.
Since $A \neq 0$ (otherwise, the solution is no meaning),
 $\therefore \delta = 0$
At $x = a$, $0 = A \sin ka$.
Since $A \neq 0$, \therefore Sinka = $0 \implies Ka = nT$ $(N = 1,2,3^{-1})$
 $\therefore K = \frac{NT}{a}$.
Here, we kick out $N = 0$ and $N < 0$ Solutions, as they have
no meaning in veality. Now, $\psi_n = A \sin (\frac{MT}{ax})$.
 $t^2 \pi^2 n^2$

 $K = \frac{n\pi}{\alpha} = \sqrt{\frac{2mE}{\hbar^2}} \implies E_n = \frac{\hbar^2 \pi^2 n^2}{2m\alpha^2}, n = 1, 2, 3, ...$

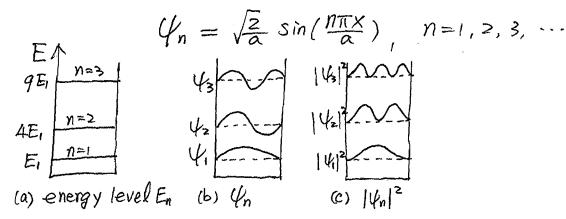
 $E_n = \frac{\hbar^2 \pi^2 n^2}{2m a^2}$ indicates that the particle energy is guantized in the infinit high potential well.

 $|\Psi(x)|^2$ is the probability density of finding the particle at position x. Since the probability of finding the particle in all space is 1 (i.e., normalization regulatement), we have $\int |\Psi(x)|^2 dx = 1$

$$\int_{0}^{A} A^{2} \sin^{2} kx \, dx = \int_{0}^{a} A^{2} \sin^{2} \frac{n\pi}{a} x \, dx = A^{2} \cdot \frac{Q}{2} = 1$$

$$\Rightarrow A = \sqrt{\frac{2}{a}}.$$

. . the normalized wave function (eigen wave function) is



Particle's motion in a potential well is a common phenomenon, e.g., the electron in hydrogen atom does 3-D motion in a the Coulomb potential, just the wall is not a square, but distributes along $-\frac{1}{r}$.

- * Note: the lowest energy E, =0, which is completely different from classical mechanics. This is due to the wave nature of particle — "a wave at rest" does not exist!
- + Note: the full wave function $(\mathcal{Y}_n(X,t) \propto Sin(\frac{n\pi X}{a})e^{-iEnt/h})$ which is a standy wave.

Example 2: Potential Barrier Renetiation (Finite Retential Well)
(leading to quantum tunnel effect) V

$$V = \begin{cases} V_0, & 0 \le x \le a \\ 0, & x < 0 \text{ and } x > a \end{cases}$$

Since $E_k = E < V_0$, according to $0 = a = x$
classical mechanics, the particle's
energy isn't high enough to overcome
the potential barrier, so it will be
reflected back. However, in quantum world, the situation
is very different: Resides reflection, the de Bros/ie above
of the particle can penetrate the potential barrier to
yeach region III.
The stationary - state Schrödiger equation is
 $-\frac{\hbar^2}{2m} \frac{d^2 \psi}{dx^2} + V(x)\psi = E \psi$
In Region I, $\frac{d^2 \psi_1}{dx^2} = -\frac{2mE}{\hbar^2} \psi_1 = -k_i^2 \psi_1, \quad k_1 = \sqrt{\frac{2mE}{\hbar}}$
Its general solution is $\psi_1 = A_1 e^{iK_1x} + B_1 e^{-eK_1x}$
Its general solution is $\psi_2 = A_2 e^{iK_2x} + B_2 e^{-CK_2x}$
Its general solution is $\psi_2 = A_2 e^{iK_2x} + B_2 e^{-CK_2x}$
In region II, the same equation as I, but no reflected wave
 $i, \quad \psi_3 = A_3 e^{iK_1x}$

The constants A, , Az, Az, B, Bz will be determined from the boundary conditions, a considering that the wave functions must be continuous at X=0 -1 X=a. We skip the procedures of determining constants, but the wave function solution is given by T II II AAI and A The particle has certain probability to occur in region III. That is, the particle has certain probability to penetrate the potential barrier. - This is the guartum tunnel effect. Applications: 1) Nuclear emission of a particle (<10 MeV) from Vadioactive nucleus potential (>20 Mev) 2) Tunnel diode invented in 1950s, for high-speed Switch or microwave device 3) Election Scanning tunneling microscope (STM) (STM technology) Can be used to Manipulate and move atoms and molecules.) (STM technology) Amplifient The Voltages added is low, but the the electron in the needle still has probabilities to penetrate the potential barrier to form 'tunneling current ". This current is very sensitive to the distance. from the needle to the surface - scanning Vey high spatial resolution) This carrent is very sensitive to the distance (Vey high spatial resolution) from the needle to the surface -> scanning -> atomic dimension) the needle gives tomography of the surface.

(2) Harmonic Oscillator
$$(1-D)$$
.
The force that a particle experiences $F = -\kappa x$,
where x is the displacement of particle velative to its
balance point 0. \therefore Epotential $= \frac{1}{2} \kappa x^2 = V$.
The stationary-state Schrödinger equation is
 $\left(-\frac{\hbar^2}{2m}\frac{\partial^2}{\partial x} + \frac{1}{2}\kappa x^2\right)\Psi = E\Psi$.
Let $\beta = \alpha x$, where $\alpha = (mk/\hbar^2)^{1/4}$.
 $\therefore \frac{d^2\Psi}{d\beta^2} + (\lambda - \beta^2)\Psi = 0$
where $\lambda = \frac{2mE}{\hbar^2 a^2} = \frac{2E}{\hbar}\sqrt{\frac{m}{K}} = \frac{2E}{\hbar w}$. $\omega = \sqrt{\kappa/m}$.
The solution to the equation is:
 $E_n = (n + \frac{1}{2})\hbar\omega$, $n=0,1,2,\cdots$.
 $\Psi_n = \left(\frac{\alpha}{\sqrt{\pi} 2^n \cdot n!}\right)^k e^{-\frac{1}{2}d^2x^2} H_n(\alpha x)$,
Where $H_n(\alpha x)$ is Hermitran polynomial
where $H_n(\alpha x)$ is Hermitran polynomial
 $Where H_n(\alpha x)$ is Hermitran polynomial
 $Where H_n(\alpha x)$ is Hermitran $polynomial$
 $M_n = \frac{1}{\sqrt{\pi} 2^n \cdot n!} = 2\alpha x$, $H_n(\alpha x) = 4(\alpha x)^2 - 2, \cdots$
 $F_{n=1} = \frac{1}{\sqrt{\pi}} e^{-\frac{1}{2}\pi 2^n - \frac{1}{\sqrt{\pi}}} = \frac{1}{\sqrt{\pi}} e^{-\frac{1}{2}\pi 2^n - \frac{1}{\sqrt{\pi}}} e^{-\frac$