

(4) Eigenvalue Equation in $\{|\vec{r}\rangle\}$ representation.

In the abstract state space, an eigenvalue equation is

$$\hat{A} |\psi\rangle = a |\psi\rangle$$

Where \hat{A} — observable operator, $|\psi\rangle$ — state vector,

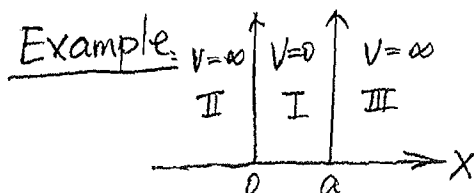
a — a complex constant (real number in reality)

It can be proven that in the $\{|\vec{r}\rangle\}$ representation, the eigenvalue equation becomes

$$\hat{A} \psi(\vec{r}) = a \psi(\vec{r})$$

Where \hat{A} is the representation of \hat{A} operator in $\{|\vec{r}\rangle\}$.

$\psi(\vec{r})$ called the ^{eigen}wave function, is the projection of the ^{eigen}state $|\psi\rangle$ on the $\{|\vec{r}\rangle\}$ representation.



(1-D) Infinite High Potential Well

A particle moves within a potential well between $x=0$ and $x=a$. The normalized wave function is given by

$$\psi_n(x) = \begin{cases} \sqrt{\frac{2}{a}} \sin\left(\frac{n\pi x}{a}\right), & 0 < x < a \\ 0, & \text{elsewhere.} \end{cases}$$

- ① Try to derive the mean of the particle's momentum and kinetic energy
- ② Verify whether this wave function is an eigenfunction of momentum, whether an eigen wavefunction of kinetic energy.

Solution: In $\{|x\rangle\}$ representation, $\hat{p} = -i\hbar \frac{\partial}{\partial x}$, $\frac{\hat{p}^2}{2m} = -\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2}$.

$$\begin{aligned} \bar{p} &= \int_{-\infty}^{+\infty} dx \psi_n^*(x) (-i\hbar \frac{\partial}{\partial x}) \psi_n(x) = -\frac{2i\hbar}{a} \int_0^a dx \sin\left(\frac{n\pi x}{a}\right) (-i\hbar \frac{\partial}{\partial x}) \sin\left(\frac{n\pi x}{a}\right) \\ &= -i\hbar \frac{2}{a} \cdot \frac{n\pi}{a} \int_0^a \sin\left(\frac{n\pi x}{a}\right) \cos\left(\frac{n\pi x}{a}\right) dx \\ &= -i\hbar \frac{2}{a} \cdot \frac{1}{2} \sin^2\left(\frac{n\pi x}{a}\right) \Big|_0^a = 0 - 0 = 0. \end{aligned}$$

Kinetic energy mean:

$$\begin{aligned} \overline{\frac{p^2}{2m}} &= \int_{-a}^{+a} \psi_n^*(x) \left(-\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} \right) \psi_n(x) dx \\ &= -\frac{\hbar^2}{2ma} \int_0^a \sin\left(\frac{n\pi x}{a}\right) \frac{\partial^2}{\partial x^2} \sin\left(\frac{n\pi x}{a}\right) dx \\ &= \frac{\hbar^2}{2ma} \left(\frac{n\pi}{a}\right)^2 \int_0^a \sin^2\left(\frac{n\pi x}{a}\right) dx \\ &= \frac{\hbar^2 \pi^2 n^2}{2ma^2} \end{aligned}$$

$$\boxed{\sin^2 \theta = \frac{1 - \cos 2\theta}{2}}$$

Verifying eigen wave function:

"Apply the operator to the wave function to see whether you can obtain a number times the same wave function."

$$(1) \hat{p} \psi_n(x) = -i\hbar \frac{\partial}{\partial x} \psi_n(x) = -i\hbar \sqrt{\frac{2}{a}} \frac{\partial}{\partial x} \sin\left(\frac{n\pi x}{a}\right) = i\sqrt{\frac{2}{a}} \frac{\hbar n\pi}{a} \cos\left(\frac{n\pi x}{a}\right)$$

— $\psi_n(x)$ is NOT \hat{p} 's eigen function.

$$(2) \frac{\hat{p}^2}{2m} \psi_n(x) = -\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} \psi_n(x) = -\frac{\hbar^2}{2m} \sqrt{\frac{2}{a}} \frac{\partial^2}{\partial x^2} \sin\left(\frac{n\pi x}{a}\right)$$

$$= \frac{\hbar^2}{2m} \sqrt{\frac{2}{a}} \left(\frac{n\pi}{a}\right)^2 \sin\left(\frac{n\pi x}{a}\right) = \frac{\hbar^2 \pi^2 n^2}{2ma^2} \psi_n(x)$$

$\therefore \psi_n(x)$ is the eigen function of kinetic energy.

The eigen value is $\frac{\hbar^2 \pi^2 n^2}{2ma^2}$, the same as the mean we derived above — of course, the mean of the kinetic energy in its own eigenstate is equal to the eigenvalue.

Thus, we solved the HW #3, Problem #3. You may practice it yourself. Then apply similar technique to HW #3, Problem #2.

(5) Schrödinger Equation in $\{|\vec{r}\rangle\}$ and $\{|\vec{p}\rangle\}$ representations.

In the abstract state space, the Schrödinger equation is

$$i\hbar \frac{d}{dt} |\psi(t)\rangle = \hat{H}(t) |\psi(t)\rangle.$$

For a (spinless) particle in a scalar potential $V(\vec{r})$, the operator

$$\hat{H} = \frac{\hat{\vec{p}}^2}{2m} + V(\hat{\vec{r}}).$$

It can be proven that in the $\{|\vec{r}\rangle\}$ representation, the Schrödinger equation becomes:

$$i\hbar \frac{\partial}{\partial t} \psi(\vec{r}, t) = \left[-\frac{\hbar^2}{2m} \nabla^2 + V(\vec{r}) \right] \psi(\vec{r}, t).$$

In the $\{|\vec{p}\rangle\}$ representation, the Schrödinger equation is

$$i\hbar \frac{\partial}{\partial t} \bar{\psi}(\vec{p}, t) = \frac{\vec{p}^2}{2m} \bar{\psi}(\vec{p}, t) + (2\pi\hbar)^{-3/2} \int d^3p' V(\vec{p}-\vec{p}') \bar{\psi}(\vec{p}', t).$$

Here, $\psi(\vec{r}, t) \equiv \langle \vec{r} | \psi \rangle$

$\bar{\psi}(\vec{p}, t) \equiv \langle \vec{p} | \psi \rangle.$

$$\bar{V}(\vec{p}) = (2\pi\hbar)^{-3/2} \int d^3r e^{-\frac{i}{\hbar} \vec{p} \cdot \vec{r}} V(\vec{r}).$$

§3.6 Solutions to Eigenvalue Equation and Schrödinger Equation

16. Solution to Eigenvalue Equation and Schrödinger Equation.

Schrödinger equation: $i\hbar \frac{\partial}{\partial t} \psi(\vec{r}, t) = \left[-\frac{\hbar^2}{2m} \nabla^2 + V(\vec{r}) \right] \psi(\vec{r}, t)$.

If $V(\vec{r})$ is not explicitly dependent on time t , then we have

$$\psi(\vec{r}, t) = \psi(\vec{r}) T(t).$$

Substituting this into the Schrödinger equation:

$$\frac{i\hbar}{T} \frac{dT}{dt} = \frac{1}{\psi(\vec{r})} \left[-\frac{\hbar^2}{2m} \nabla^2 + V(\vec{r}) \right] \psi(\vec{r}) = E$$

↑
total energy

$$\therefore \frac{i\hbar}{T} \frac{dT}{dt} = E \Rightarrow T = T_0 e^{-iEt/\hbar}$$

$$\left\{ \begin{array}{l} \left[-\frac{\hbar^2}{2m} \nabla^2 + V(\vec{r}) \right] \psi(\vec{r}) = E \psi(\vec{r}) \end{array} \right.$$

$$\therefore \psi(\vec{r}, t) = \psi(\vec{r}) e^{-iEt/\hbar}$$

Probability density $= |\psi(\vec{r}, t)|^2 = |\psi(\vec{r})|^2$ is independent of t .
i.e., the probability of the particle appearing at position \vec{r} does not change with time!

Equation $\left[-\frac{\hbar^2}{2m} \nabla^2 + V(\vec{r}) \right] \psi(\vec{r}) = E \psi(\vec{r})$

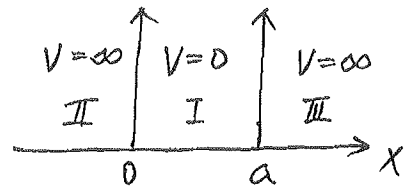
is called the stationary state Schrödinger equation. Essentially, it is the energy eigenvalue equation.

Here, we show a few examples of how to solve the stationary state Schrödinger equation, i.e., the energy eigenvalue equation, to derive the system states and eigenvalues.

(1) 1-Dimension Infinite potential well

The stationary-state Schrödinger equation

$$\text{is: } \left[-\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} + V \right] \psi(x) = E \psi(x).$$



$$V(x) = \begin{cases} \infty, & x < 0 \\ 0, & 0 < x < a \\ \infty, & x > a \end{cases}$$

Since a particle cannot be in an infinite potential, $\therefore \psi = 0$ in regions II and III. (Physics)

In region I, $V = 0$, the equation is simplified to

$$-\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} \psi(x) = E \psi(x)$$

Let $k \equiv \sqrt{\frac{2mE}{\hbar^2}}$, then: $\frac{d^2 \psi}{dx^2} + k^2 \psi = 0$

The general solution to this equation is

$$\psi = A \sin(kx + \delta),$$

where A and δ are constants to be determined from boundary conditions and normalization requirements.

Considering from physics aspects, since the particle cannot be in the $V = \infty$ region, i.e., the probability to be in regions II and III is zero. Therefore, $\psi(x=0) = 0$, $\psi(x=a) = 0$.

At $x=0$, $0 = A \sin \delta$.

Since $A \neq 0$ (otherwise, the solution is no meaning),

$$\therefore \delta = 0$$

At $x=a$, $0 = A \sin ka$.

Since $A \neq 0$, $\therefore \sin ka = 0 \Rightarrow ka = n\pi$ ($n = 1, 2, 3, \dots$)

$$\therefore k = \frac{n\pi}{a}$$

Here, we kick out $n=0$ and $n < 0$ solutions, as they have no meaning in reality. Now: $\psi_n = A \sin\left(\frac{n\pi}{a}x\right)$.

$$\therefore k = \frac{n\pi}{a} = \sqrt{\frac{2mE}{\hbar^2}} \Rightarrow E_n = \frac{\hbar^2 \pi^2 n^2}{2ma^2}, \quad n = 1, 2, 3, \dots$$

$E_n = \frac{\hbar^2 \pi^2 n^2}{2ma^2}$ indicates that the particle energy is quantized in the infinite high potential well.

$|\psi(x)|^2$ is the probability density of finding the particle at position x . Since the probability of finding the particle in all space is 1 (i.e., normalization requirement), we have

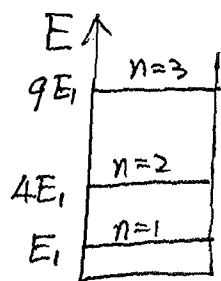
$$\int |\psi(x)|^2 dx = 1$$

$$\therefore \int_0^a A^2 \sin^2 kx dx = \int_0^a A^2 \sin^2 \frac{n\pi}{a} x dx = A^2 \cdot \frac{a}{2} = 1$$

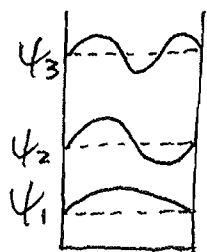
$$\Rightarrow A = \sqrt{\frac{2}{a}}$$

\therefore the normalized wave function (eigen wave function) is

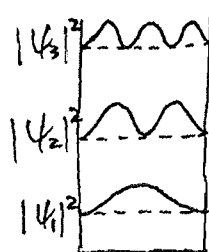
$$\psi_n = \sqrt{\frac{2}{a}} \sin\left(\frac{n\pi x}{a}\right), \quad n=1, 2, 3, \dots$$



(a) energy level E_n



(b) ψ_n



(c) $|\psi_n|^2$

Particle's motion in a potential well is a common phenomenon, e.g., the electron in hydrogen atom does 3-D motion in the Coulomb potential, just the wall is not a square, but distributes along $-\frac{1}{r}$.

* Note: the lowest energy $E_1 \neq 0$, which is completely different from classical mechanics. This is due to the wave nature of particle — "a wave at rest" does not exist!

* Note: the full wave function $\psi_n(x,t) \propto \sin\left(\frac{n\pi x}{a}\right) e^{-iE_n t/\hbar}$, which is a standing wave.

(2) Harmonic Oscillator (1-D):

The force that a particle experiences $F = -kx$,
 where x is the displacement of particle relative to its
 balance point 0. $\therefore E_{\text{potential}} = \frac{1}{2} kx^2 = V$.

The stationary-state Schrödinger equation is

$$\left(-\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} + \frac{1}{2} kx^2 \right) \psi = E\psi$$

Let $\beta = \alpha x$, where $\alpha = (mk/\hbar^2)^{1/4}$.

$$\therefore \frac{d^2 \psi}{d\beta^2} + (\lambda - \beta^2) \psi = 0$$

where $\lambda = \frac{2mE}{\hbar^2 \alpha^2} = \frac{2E}{\hbar} \sqrt{\frac{m}{k}} = \frac{2E}{\hbar \omega}$. $\omega = \sqrt{k/m}$.

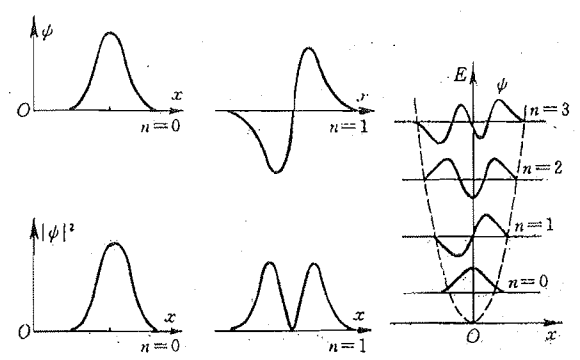
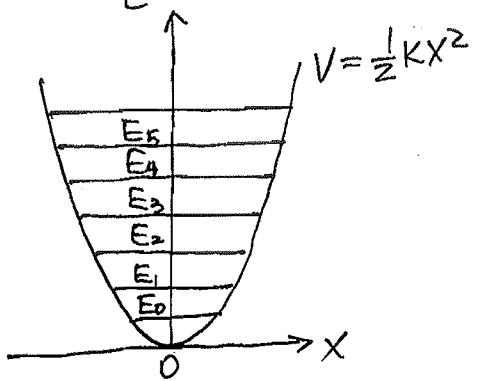
The solution to the equation is:

$$E_n = \left(n + \frac{1}{2} \right) \hbar \omega, \quad n=0, 1, 2, \dots$$

$$\psi_n = \left(\frac{\alpha}{\sqrt{\pi} 2^n n!} \right)^{1/2} e^{-\frac{1}{2} \alpha^2 x^2} H_n(\alpha x)$$

Where $H_n(\alpha x)$ is Hermitean polynomial

$$H_0(\alpha x) = 1, \quad H_1(\alpha x) = 2\alpha x, \quad H_2(\alpha x) = 4(\alpha x)^2 - 2, \dots$$



① Plot energy levels within the potential energy curve. The length of horizontal lines shows the oscillator motion range.

② When $\hbar = 0$,
 $E_0 = \frac{1}{2} \hbar \omega \neq 0$
 \Rightarrow no oscillator at rest!

③ E_n are equally separated — similar to Planck's hypothesis — quanta of oscillator energy!